

# Quantum Gravity meets Statistical Physics II

Alex Belin - CERN

Oct 26<sup>th</sup> @ Banff Intl Research Station

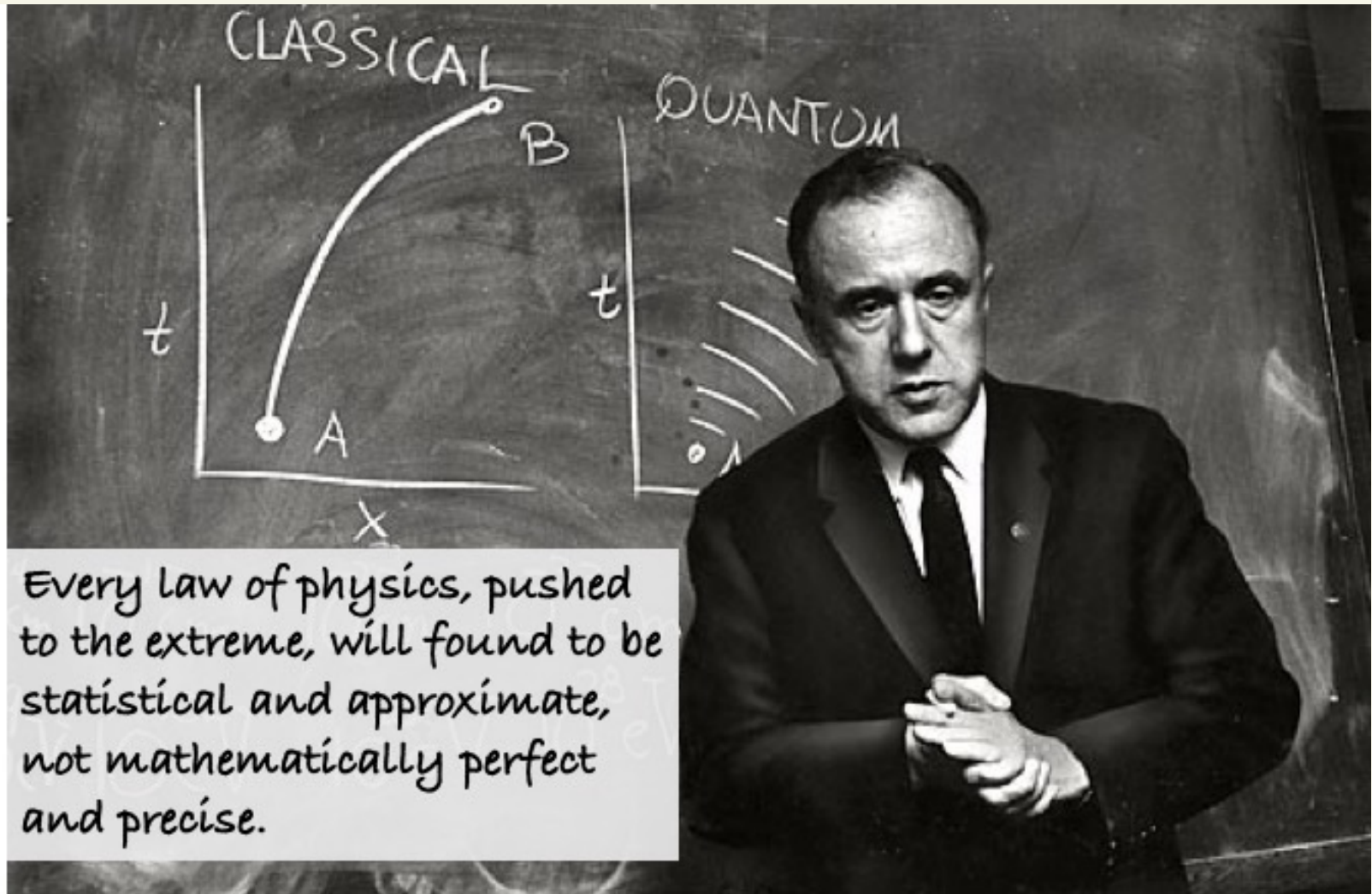
Based on:

2006.05499 w/ J. de Boer

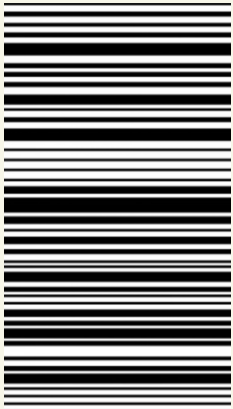
2110.xxxxx w/ J. de Boer, D. Listka

2111.xxxxx w/ J. de Boer, P. Nayak, J. Sonner

2111.xxxxx w/ T. Anous, J. de Boer, D. Listka



⇒ This seems like the right mindset to think about semi-classical gravity!

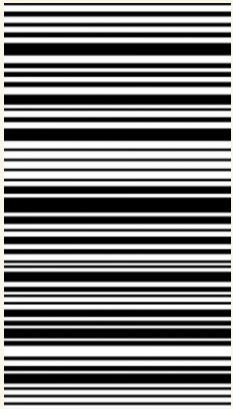


$E_n$

$$R_{mn} = \langle E_m | O | E_n \rangle$$

Microscopies

0.78	0.2	0.85	-0.24	-0.98	-0.74	-0.46	-0.86	0.52	0.29	-0.25	1.	0.29	0.9	0.93	0.66	0.95	-0.34	-0.01	0.67
0.05	0.23	-0.98	0.84	-0.86	-0.88	-0.15	0.22	-0.8	0.42	-0.94	-0.34	-0.85	-0.31	-0.22	-0.92	0.1	-0.58	-0.65	-0.44
-0.23	-0.48	-0.82	0.07	-0.88	0.52	0.81	0.1	-0.23	-0.87	-0.89	0.69	0.19	-0.15	-0.73	-0.52	-0.23	0.64	-0.7	-0.56
0.94	0.1	-0.78	0.5	1.	0.98	0.03	-0.86	0.66	0.75	0.11	0.85	0.94	0.19	-0.78	-0.38	-0.41	-0.84	0.91	0.94
-0.78	0.67	0.58	0.58	-0.3	-0.16	-0.75	0.84	0.48	0.67	-0.31	0.89	0.43	0.26	0.26	0.81	-0.53	0.12	-0.58	0.53
-0.32	-0.49	-0.29	0.35	0.18	-0.55	-0.75	0.2	-0.81	-0.14	-0.58	-0.21	-0.88	0.69	0.46	0.82	-0.26	-0.77	0.92	0.6
0.48	0.84	-0.41	0.32	-0.84	-0.58	0.8	0.84	-0.75	-0.52	-0.95	-0.52	0.81	0.49	-0.57	0.21	0.4	0.77	-0.85	-0.14
0.29	-0.04	-0.05	0.64	-0.68	0.06	0.3	0.88	-0.78	-0.72	0.77	-0.57	-0.54	0.35	-0.68	-0.86	0.48	0.6	0.76	-0.86
-0.67	-0.53	-0.23	-0.47	0.21	0.65	0.29	-0.65	-0.66	-0.76	-0.16	0.79	-0.21	0.83	0.46	0.65	-0.38	-0.37	0.5	-0.86
0.5	-0.82	0.07	0.12	0.83	0.12	-0.83	0.78	0.24	-0.16	0.36	-0.76	0.95	-0.95	0.71	-0.38	0.77	0.92	0.57	0.67
0.08	0.14	0.37	0.64	-0.82	0.51	0.78	-0.81	-0.85	-0.46	-0.97	-0.13	0.46	0.29	0.	0.9	-0.91	-0.56	-0.42	0.26
-0.96	0.9	-0.47	0.01	-0.94	-0.52	-0.77	-0.62	0.62	-0.14	0.98	0.85	-0.3	-0.88	-0.01	-0.69	-0.24	-0.7	-0.39	-0.57
0.82	-0.18	0.58	-0.83	-0.34	-0.89	-0.81	0.97	0.54	-0.27	-0.23	-0.76	-0.84	-0.49	0.11	0.41	-0.37	-0.24	-0.31	0.99
-0.68	0.69	-0.88	0.16	0.98	0.75	-0.94	0.27	0.41	-0.99	-0.27	0.38	-0.59	0.8	0.43	-0.63	-0.4	-0.57	-0.14	0.87
-0.12	-0.51	-0.17	0.61	0.87	0.74	0.35	-0.48	-0.42	0.	-0.86	-0.47	-0.53	-0.99	-0.37	-0.16	0.58	-0.95	-0.82	0.8
0.56	-0.66	-0.99	-0.66	0.57	-0.18	0.92	0.14	-0.27	0.73	-0.36	-0.45	0.84	-0.27	-0.46	-0.75	0.62	-0.89	0.39	0.14
0.07	0.04	0.98	-0.43	-0.73	0.97	0.56	0.19	0.63	-0.24	0.22	0.63	-0.75	0.05	0.34	0.6	-0.04	-0.15	0.72	0.42
-0.45	-0.9	0.84	-0.99	0.8	0.78	0.72	0.94	0.15	-0.46	0.64	-0.34	-0.79	1.	-0.47	0.53	-0.38	-0.81	0.69	0.89
-0.54	-0.83	-0.57	0.71	0.23	-0.15	-0.85	-0.93	-0.7	0.43	-0.98	0.76	-0.95	0.86	-0.35	0.99	0.91	-0.19	-0.29	0.76
-0.78	-0.86	0.85	-0.69	0.	-0.98	0.35	0.12	-0.67	0.38	-0.69	0.36	-0.82	-0.34	-0.79	0.73	0.07	0.99	-0.07	-0.69



$E_n$

$$R_{mn} = \langle E_m | O | E_n \rangle$$

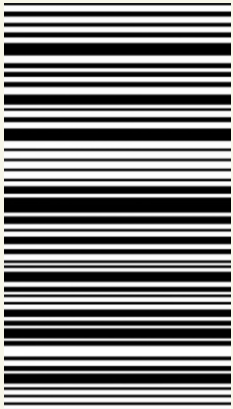
Microscopies

0.78	0.2	0.85	-0.24	-0.98	-0.74	-0.46	-0.86	0.52	0.29	-0.25	1.	0.29	0.9	0.93	0.66	0.95	-0.34	-0.01	0.67
0.05	0.23	-0.98	0.84	-0.86	-0.88	-0.15	0.22	-0.8	0.42	-0.94	-0.34	-0.85	-0.31	-0.22	-0.92	0.1	-0.58	-0.65	-0.44
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-0.78	0.67	0.58	0.58	-0.3	-0.16	-0.75	0.84	0.48	0.67	-0.31	0.89	0.43	0.26	0.26	0.81	-0.59	0.12	-0.58	0.53
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-0.67	-0.53	-0.23	-0.47	0.21	0.65	0.29	-0.65	-0.66	-0.76	0.16	0.79	-0.21	0.83	0.46	0.65	-0.38	-0.37	0.5	-0.86
0.5	-0.82	0.07	0.12	0.83	0.12	-0.93	0.78	0.24	-0.16	0.36	-0.76	0.95	-0.95	0.71	-0.38	0.77	0.92	0.57	0.67
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-0.12	-0.51	-0.17	0.61	0.87	0.74	0.35	-0.48	-0.42	0.	-0.86	-0.47	-0.53	-0.99	-0.37	-0.16	0.58	-0.95	-0.82	0.8
0.56	-0.66	-0.99	-0.66	0.57	-0.18	0.92	0.14	-0.27	0.73	-0.36	-0.45	0.84	-0.27	-0.46	-0.75	0.62	-0.89	0.39	0.14
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-0.78	-0.86	0.85	-0.69	0.	-0.98	0.35	0.12	-0.67	0.38	-0.69	0.36	-0.82	-0.34	-0.79	0.73	0.07	0.99	-0.07	-0.69

Semi-classical gravity

=

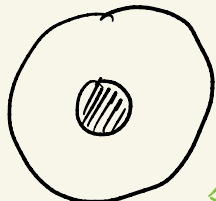
The thg of the stat. distribution of  
 $\{E_n, R_{mn}\}$



$$E_n \quad R_{mn} = \langle E_m | O | E_n \rangle$$

Microscopies

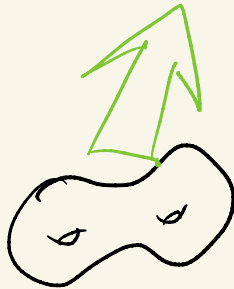
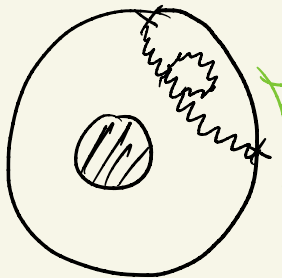
0.78	0.2	0.85	-0.24	-0.98	-0.74	-0.46	-0.86	0.52	0.29	-0.25	1.	0.29	0.9	0.93	0.66	0.95	-0.34	-0.01	0.07
0.05	0.23	-0.98	0.84	-0.86	-0.88	-0.15	0.22	-0.8	0.42	-0.94	-0.34	-0.85	-0.31	-0.22	-0.92	0.1	-0.58	-0.65	-0.44
-0.23	-0.48	-0.82	0.07	-0.88	0.52	0.81	0.1	-0.23	-0.87	-0.89	0.69	0.19	-0.15	-0.73	-0.52	-0.23	0.64	-0.7	-0.66
0.94	0.1	-0.78	0.5	1.	0.98	0.03	-0.86	0.66	0.75	0.11	0.85	0.94	0.19	-0.78	-0.38	-0.41	-0.84	0.91	0.94
-0.78	0.67	0.58	0.58	-0.3	-0.16	-0.75	0.84	-0.48	0.67	-0.31	0.89	0.43	0.26	0.26	0.81	-0.53	0.12	-0.58	0.53
-0.32	-0.49	-0.29	0.35	0.18	-0.55	-0.75	0.2	-0.81	-0.14	-0.58	-0.21	-0.88	0.69	0.46	0.82	-0.26	-0.77	0.92	0.6
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0.29	-0.04	-0.05	0.64	-0.68	0.06	0.3	0.88	-0.78	-0.72	0.77	-0.57	-0.54	0.35	-0.68	-0.86	0.48	0.6	0.76	-0.86
-0.67	-0.53	-0.23	-0.47	0.21	0.65	0.29	-0.65	-0.66	-0.76	0.16	0.79	-0.21	0.83	0.46	0.65	-0.38	-0.37	0.5	-0.86
0.5	-0.82	0.07	0.12	0.83	0.12	-0.03	0.78	0.24	-0.16	0.36	-0.76	0.95	-0.95	0.71	-0.38	0.77	0.92	0.57	0.67
0.08	0.14	0.37	0.64	-0.82	0.51	0.78	-0.81	-0.85	-0.45	-0.97	-0.13	0.45	0.29	0.	0.9	-0.91	-0.56	-0.42	0.26
-0.96	0.9	-0.47	0.01	-0.94	-0.52	-0.77	-0.62	0.62	-0.14	0.98	0.85	-0.3	-0.88	-0.01	-0.69	-0.24	-0.7	-0.39	-0.57
0.82	-0.18	0.58	-0.83	-0.34	-0.89	-0.81	0.97	0.54	-0.27	-0.23	-0.76	-0.84	-0.49	0.11	0.41	-0.37	-0.24	-0.31	0.99
-0.68	0.69	-0.88	0.16	0.98	0.75	-0.94	0.27	0.41	-0.99	-0.27	0.38	-0.59	0.8	0.43	-0.63	-0.4	-0.57	-0.14	0.87
-0.12	-0.51	-0.17	0.61	0.87	0.74	0.35	-0.48	-0.42	0.	-0.86	-0.47	-0.53	-0.99	-0.37	-0.16	0.58	-0.95	-0.82	0.8
0.56	-0.66	-0.99	-0.66	0.57	-0.18	0.92	0.14	-0.27	0.73	-0.36	-0.45	0.84	-0.27	-0.46	-0.75	0.62	-0.89	0.39	0.14
0.07	0.04	0.98	-0.43	-0.73	0.97	0.56	0.19	0.63	-0.24	0.22	0.63	-0.75	0.05	0.34	0.6	-0.04	-0.15	0.72	0.42
-0.45	-0.9	0.84	-0.99	0.8	0.78	0.72	0.94	0.15	-0.46	0.64	-0.34	-0.79	1.	-0.47	0.53	-0.38	-0.81	0.69	0.89
-0.54	-0.83	-0.57	0.71	0.23	-0.15	-0.85	-0.93	-0.7	0.43	-0.98	0.76	-0.95	0.86	-0.35	0.99	0.91	-0.19	-0.29	0.76
-0.78	-0.86	0.85	-0.69	0.	-0.98	0.35	0.12	-0.67	0.38	-0.69	0.36	-0.82	-0.34	-0.79	0.73	0.07	0.99	-0.07	-0.69

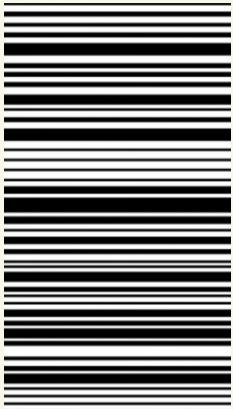


$$S = \frac{A}{4G\hbar}$$



Semi-classical gravity  
=  
The thg of the stat. distribution of  
 $\{E_n, R_{mn}\}$



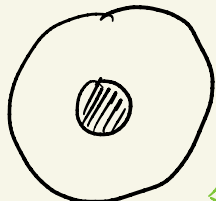


$E_n$

$$R_{mn} = \langle E_m | O | E_n \rangle$$

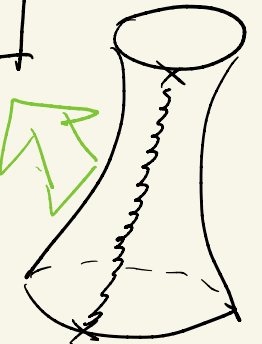
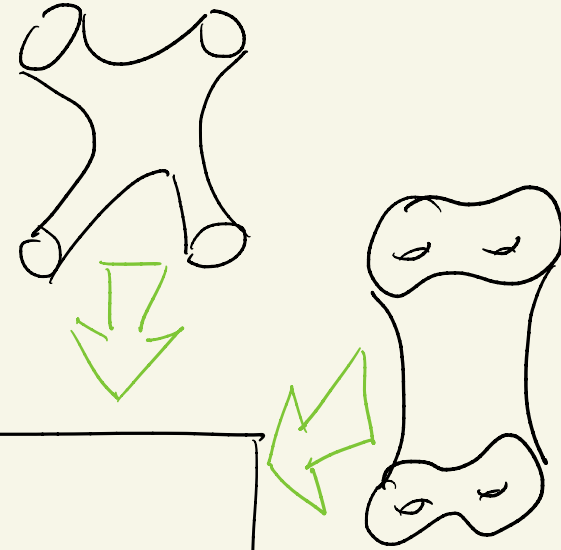
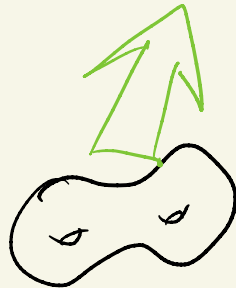
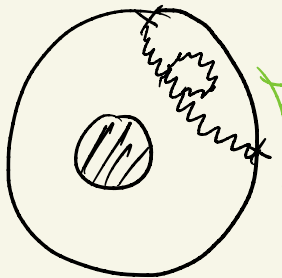
Microscopies

0.78	0.2	0.85	-0.24	-0.98	-0.74	-0.46	-0.86	0.52	0.29	-0.25	1.	0.29	0.9	0.93	0.66	0.95	-0.34	-0.01	0.97
0.05	0.23	-0.98	0.84	-0.86	-0.88	-0.15	0.22	-0.8	0.42	-0.94	-0.34	-0.85	-0.31	-0.22	-0.92	0.1	-0.58	-0.65	-0.44
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0.94	0.1	-0.78	0.5	1.	0.98	0.03	-0.86	0.66	0.75	0.11	0.85	0.94	0.19	-0.78	-0.38	-0.41	-0.84	0.91	0.94
-0.78	0.67	0.58	0.58	-0.3	-0.16	-0.75	0.84	-0.48	0.67	-0.31	0.89	0.43	0.26	0.26	0.81	-0.59	0.12	-0.58	0.53
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0.29	-0.04	-0.05	0.64	-0.68	0.06	0.3	0.88	-0.78	-0.72	0.77	-0.57	-0.54	0.35	-0.68	-0.86	0.48	0.6	0.76	-0.86
-0.67	-0.53	-0.23	-0.47	0.21	0.65	0.29	-0.65	-0.66	-0.76	0.16	0.79	-0.21	0.83	0.46	0.65	-0.38	-0.37	0.5	-0.86
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-0.96	0.9	-0.47	0.01	-0.94	-0.52	-0.77	-0.62	0.62	0.14	0.98	0.85	-0.3	-0.88	-0.01	-0.69	-0.24	-0.7	-0.39	-0.57
0.82	-0.18	0.58	-0.83	-0.34	-0.89	-0.81	0.97	0.54	-0.27	-0.23	-0.76	-0.84	-0.49	0.11	0.41	-0.37	-0.24	-0.31	0.99
-0.68	0.69	-0.88	0.16	0.98	0.75	-0.84	0.27	0.41	-0.99	-0.27	0.38	-0.59	0.8	0.43	-0.63	-0.4	-0.57	-0.14	0.87
-0.12	-0.51	-0.17	0.61	0.87	0.74	0.35	-0.48	-0.42	0.	-0.86	-0.47	-0.53	-0.99	-0.37	-0.16	0.58	-0.95	-0.82	0.8
0.56	-0.66	-0.99	-0.66	0.57	-0.18	0.92	0.14	-0.27	0.73	-0.36	-0.45	0.84	-0.27	-0.46	-0.75	0.62	-0.89	0.39	0.14
0.07	0.04	0.98	-0.43	-0.73	0.07	0.16	0.19	0.63	-0.24	0.22	0.63	-0.75	0.05	0.34	0.6	-0.04	-0.15	0.72	0.42
-0.45	-0.9	0.84	-0.99	0.8	0.78	0.72	0.94	0.15	-0.46	0.64	-0.34	-0.79	1.	-0.47	0.53	-0.38	-0.81	0.69	0.89
-0.54	-0.83	-0.57	0.71	0.23	-0.15	-0.85	-0.93	-0.7	0.43	-0.98	0.76	-0.95	0.86	-0.35	0.99	0.91	-0.19	-0.29	0.76
-0.78	-0.86	0.85	-0.69	0.	-0.98	0.35	0.12	-0.67	0.38	-0.69	0.36	-0.82	-0.34	-0.79	0.73	0.07	0.99	-0.07	-0.69



$$S = \frac{A}{4G_N}$$

Semi-classical gravity  
=  
The thg of the stat. distribution of  
 $\{E_n, R_{mn}\}$



Today → Focus on operator statistics

The ETH Ansatz:

$$\langle E_m | O | E_n \rangle = \delta_{mn} f_0(\bar{E}) + R_{mn} e^{-S(\bar{E})/2} g_0(\bar{E}, \delta E)$$

- $f_0, g_0 \rightarrow$  smooth fcts, microcan. 1 and 2 pt fcts
- $R_{mn} \rightarrow$  Indep. random numbers, 0 mean, 1 variance

Today  $\rightarrow$  Focus on operator statistics

The ETH Ansatz:

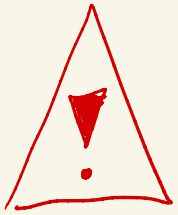
$$\langle E_m | O | E_n \rangle = \delta_{mn} f_0(\bar{E}) + R_{mn} e^{-S(\bar{E})/2} g_0(\bar{E}, \delta E)$$

- $f_0, g_0 \rightarrow$  smooth fcts, microcan. 1 and 2 pt fcts
- $R_{mn} \rightarrow$  Indep. "pseudorandom" numbers, 0 mean, 1 variance



$R_{mn}$  not actually random, fixed in any theory





$R_{mn}$  are neither indep. nor Gaussian

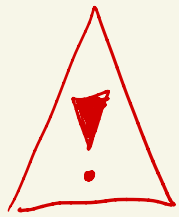


$\Rightarrow$  Inconsistent with higher point functions

Need non-trivial

$$\overline{R_{ij} R_{jk} R_{kl} R_{li}}$$

(= quartic moment with cyclic struct.)



$R_{mn}$  are neither indep. nor Gaussian



$\Rightarrow$  Inconsistent with higher point functions

Need non-trivial

$$\overline{R_{ij} R_{jk} R_{kl} R_{li}}$$

(= quartic moment with cyclic struct.)

However :

$$R_{mn} | k\text{-th moment} \sim e^{-\frac{k-1}{k} S}$$

[Foini, Kurchan]

The important point : Higher moments are suppressed

ETH . QM  $\rightarrow$  CFT

$$\langle E_m | O | E_n \rangle = \langle O_m O O_n \rangle = C_{nmo}$$

$$\Delta_{n,m} \rightarrow \infty$$

$$\Delta_o \text{ fixed}$$



ETH only applies to primaries

But in a CFT, doesn't capture all the dynamics

$$C_{ijz} \quad \Delta_{i,j,z} \rightarrow \infty$$

$$C_{00i} \quad \Delta_i \rightarrow \infty$$

} ??  
..

## New observables / probes of chaos

$$Z_{g=2}(\beta) = \sum_{0_1, 0_2, 0_3} |C_{123}|^2 e^{-\beta(\Delta_1 + \Delta_2 + \Delta_3)}$$

The genus-2 SFF:

$$F(t) = Z_{g=2}(\beta + it) Z_{g=2}(\beta - it)$$

ETH / RMT  $\Rightarrow$  doesn't offer much insight here...

# ORH conjecture

[AB, de Boer]

In a chaotic CFT:

$$C_{00i} = f_i(\bar{E}) R_i$$

$$C_{ijk} = f_i(\bar{E}, \delta E_i) R_{ijk}$$

pseudorandom variables  
0 mean, 1 variance

## Gravitational implications

Semi-classical gravity has access to the smooth facts.

But no more  $\Rightarrow$  explains lack of factorization

# Plan

① Some evidence for ORH

② Wormhole implications

# How to check ORH?

Numerically  $\rightarrow$  Way beyond current bootstrap

Consistency checks  $\rightarrow$  constraints from cross + mod. inv.

$$\begin{array}{c} \circ \\ \diagup \\ \text{---} \parallel \text{---} \\ \diagdown \\ \circ \end{array} \quad \begin{array}{c} \circ \\ \diagdown \\ \text{---} \\ \diagup \\ \circ \end{array} = \sum_H \begin{array}{c} \circ \quad \circ \\ \diagdown \quad \diagup \\ \text{---} \text{H} \\ \diagup \quad \diagdown \\ \circ \quad \circ \end{array} \Rightarrow \sum_H |C_{00H}|^2$$

[Pappadulo, Rychkov, Espin, Paltazzi]

$\Rightarrow$  Asymptotic formulas

[many people]

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[Pappadulo, Rychkov, Espin, Rattazzi]

$\Rightarrow$  Asymptotic formulas

[many people]

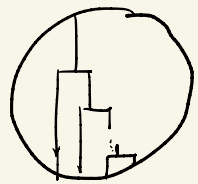
$$\sum_{i,j,k} |C_{ijk}|^2 \approx \left(\frac{27}{16}\right)^{3\Delta} e^{-3\pi\sqrt{\frac{c}{3}}\Delta}$$

[Cardy, Maloney, Maxfield]  
[Collier, Maloney, Maxfield, Tsiaras]

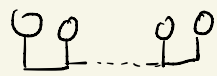


# Higher Moments

Trick: exploit modular invariance at genus- $g$



$$C_{ijk} \mid \begin{array}{l} k\text{-th moment} \\ \text{skyline} \end{array} \sim e^{-\frac{5k-4}{4k} S}$$



$$C_{ijk} \mid \begin{array}{l} k\text{-th moment} \\ \text{comb} \end{array} \sim e^{-\frac{9k-6}{8k} S}$$

[AB, de Boer, Liska - to appear]

Can also extract higher moments of

$$C_{LLH}, \quad C_{HHH} \quad \text{in } d > 2$$

[Anous, AB, de Boer, Liska - to appear]

# A generating function

$$Z[J_{abc}] = \exp \left[ f_1 J_{abc} J^{abc} + f_2 J_{aab} J^{bcc} + \sum_i g_i \text{CCCC} | i\text{-type contraction} + \dots \right]$$

$$\overline{C \dots C} = \frac{\delta}{\delta J} \dots \frac{\delta}{\delta J} Z \Big|_{J=0}$$

The functions  $f_i, g_i, \dots$  can be extracted

from asymptotic formulas

# ORH from RMT?

We can derive a form of ETH from RMT

$$\mathbb{O} = \sum_{\alpha} \lambda_{\alpha} |\alpha\rangle\langle\alpha|$$

$$\langle\alpha|E_n\rangle = U_{\alpha n}$$

↳ Haar-Random

[Atland, Bagrets, Nagak, Sonner, Vielma]

$\mathbb{O}$ : linear operator on  $\mathcal{H}^{\otimes 3}$

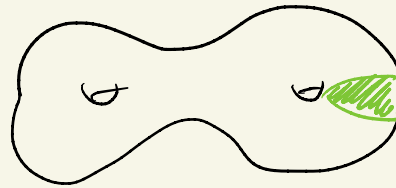
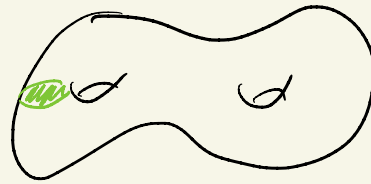
$$\langle ijkl | \mathbb{O} | lmn \rangle = C_{ijk} C_{lmn}$$

Assume  $\mathbb{O}$  is Haar-Random  $\Rightarrow$  derive ORH

[AB, de Boer, Nagak, Sonner - to appear]

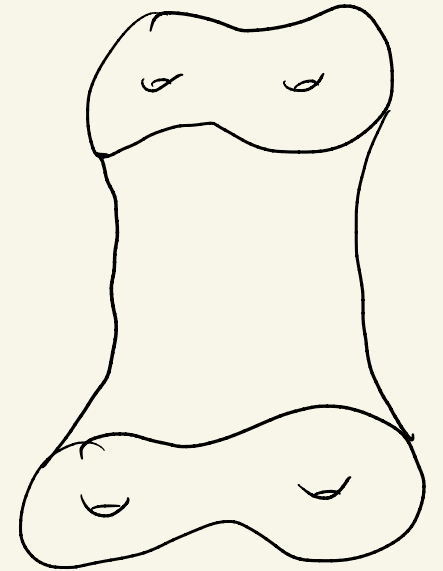
# Implications for gravity

$$Z_{g=2} \times Z_{g=2} \Big|_{\text{grav}} =$$



Disconnected  
Handlebodies

+



Connected  
genus-2 wormhole

$\approx$

$$e^{\frac{c}{2} \frac{\pi^2}{\beta}}$$

+



vanishing on-shell  
action

$$Z_{g=2} \times Z_{g=2} \Big|_{\text{ORH}} \approx \sum_{\Delta_i} \underbrace{C C' C C'}_{\text{red}} e^{-6\beta\Delta}$$

$$\underbrace{e^{\frac{c}{2} \frac{\pi^2}{\beta}}}_{\text{green}} + \frac{\sum_{\Delta} e^{-6\beta\Delta}}{\text{red}}$$

↳ No. large saddle  
 $\approx 1$

Add quartic moments:

$$\sum_{\Delta} C C C C |_{\text{skyline}} e^{-6\beta\Delta} = \sum_{\Delta} e^{-6\beta\Delta}$$

⇒ Same order! To check if 1-loop determinants match

need genus-2 + genus-3 handlebodies

A new connected saddle?

$$Z_{\text{dumbbell}} = \sum C_{ij} C_{jkk} e^{-3\beta\Delta}$$

$$\left( Z_{\text{dumbbell}} \right)^2 \Big|_{\text{Comb}} \sim e^{\frac{25c - 360\Delta\chi}{288} \frac{\pi^2}{\beta}}$$

If  $\Delta\chi$  small enough  $\rightarrow$  New large connected saddle, bigger than genus-2 wormhole!

A new connected saddle?

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If  $\Delta\chi$  small enough  $\rightarrow$  New large connected saddle, bigger than genus-2 wormhole!

$$\Delta\chi < \frac{c-1}{12} \left( 1 - \frac{1}{6^2} \right) \Rightarrow \text{weight of } \mathbb{Z}_6 \text{ conical defect.}$$

# Open Questions

- ORH predicts new wormhole solution, why haven't we found it? Matter supported?

- There seems to be 2 possibilities:  
Typicality / Haar Averages + Asymptotic formulas  $\Rightarrow$  predicts wormholes

or

Wormholes = everything you cannot predict from Typicality

Multi-trace terms in  $Z[J]$ ?



Thank You !

