> Significance In the On-Off Problem

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## What Can HEP and Astrophysics Practice Teach Each Other?

Astrophysics (especially $\gamma$ ray) aims at simple formulae (very fast) calculates o's directly (Asymptotic Normal) hope it's a good formula
HEP
(especially Fermilab practice)
calculates probabilities by MC (general; slow) translates into $\sigma$ 's for communication loses track of analytic structure

Cousins, Linnemann, Tucker, NIM A 595 (2008) 480-501

## Observed vs. Prospective

## Significance

- This discussion: Observed Significance (of my data)
- Post-hoc: (after data)
- Need definition of Z
- Choice of Zmin for observational claim $=\max \mathrm{P}$ (observed|background)
- Prospective Observability (before data, to optimize expt.)



## Backgrounds in Astro and HEP

- Astrophysics: On-source vs. Off-source
- side observation with $\tau=$ Toff/Ton (sensitivity ratio)

$$
\hat{\mu}_{b}=n_{\text {off }} / \tau ;
$$

- HEP: estimate background in
defined sideband region
$\tau$ is ratio of signal and sideband region


## Li and Ma (Gamma Ray)

$$
\begin{gathered}
Z_{P L}=\sqrt{2} \sqrt{x \bullet \operatorname{Ln}\left[(1+\tau)\left(\frac{x}{x+y}\right)\right]+y \bullet \operatorname{Ln}\left[\left(\frac{1+\tau}{\tau}\right)\left(\frac{y}{x+y}\right)\right]} \\
x=\mathrm{Non} ; \mathrm{y}=\mathrm{Noff}
\end{gathered}
$$

Generic test for composite hypothesis

+ Wilks' Theorem (conditions not satisfied)


## Binomial Proportion Test: Ratio of Poisson Means

P-value $=\operatorname{Pr} \operatorname{Binomial}(\geq \operatorname{Non} \mid \mathrm{p}, \mathrm{k})$ where $\mathrm{p}=\alpha /(1+\alpha)$

$$
p-\text { value }=\sum_{j==}^{k}\left(\begin{array}{l}
k \\
j
\end{array} p^{j}(1-p)^{k-j}\right.
$$

Holds $\mathrm{k}=$ Non + Noff fixed ( k a nuisance parameter)
UMPU (Uniformly Most Powerful Unbiased)
for Composite Hypothesis test $\mu_{\text {on }} / \alpha \mu_{\text {off }}>1$
Optimal? Not continuous-issues for small n
Not in common use; probably should be
Known in HEP and Astrophysics: but not as optimal, nor standard procedure

- Zhang and Ramsden claim too conservative for $Z$ small Even if true, we want $Z>4$
- Closed form in term of special functions, or sums
- Applying for large N requires some delicacy; ZpL


## Bayesian Methods

- In common use in HEP
- Cousins \& Highland "smeared likelihood" efficiency
- Predictive Posterior (after background measurement)
$\mathrm{P}($ Non $\mid$ Noff $) \quad$ (integrate posterior $\mu_{\mathrm{b}}$ )
A flat prior for background, gives Gamma dist. for $\mathrm{p}\left(\mu_{\mathrm{b}} \mid\right.$ Noff $)$

P value calc using Gamma: (also Alexandreas--Astro)
IDENTICAL to Frequentist Binomial Test

## Predictive Posterior Bayes P-value (HEP)

In words: tail sum averaged over Bayes posterior for mean or: integrate before sum

$$
P-\operatorname{value}(x, y)=\sum_{j=x}^{\infty} p(j \mid y)
$$

$$
p(j \mid y)=\int p(j \mid \mu) p(\mu \mid y) d \mu
$$

$$
p(j \mid \mu)=\frac{\mu^{j} e^{-\mu}}{j!}
$$

$$
p_{\Gamma}(\mu \mid y)=\frac{\beta^{y} e^{-\beta}}{y!}, \quad \beta=\mu / \alpha
$$

$$
p_{N}(\mu \mid y)=\operatorname{Normal}[(\mu-b) / \delta b]
$$

## Comparing the Methods

Some test cases from published literature
And a few artificial cases
Range of Non, Noff values
Different $\tau$ values (mostly $>1$ )

Can show some approximate Z's strictly > others including popular shortcut formulas
Others cross over as $\tau$ varies

Coverage Calculations (Tucker, Cousins)

| Reference | [40] | [41] | [42] | [43] | [44] | [44] | [45] | [46] |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $n_{\text {on }}$ | 4 | 6 | 9 | 17 | 50 | 67 | 200 | 523 |
| $n_{\text {off }}$ | 5 | 18.78 | 17.83 | 40.11 | 55 | 15 | 10 | 2327 |
| $\tau$ | 5.0 | 14.44 | 4.69 | 10.56 | 2.0 | 0.5 | 0.1 | 5.99 |
| $\hat{\mu}_{\text {b }}$ | 1.0 | 1.3 | 3.8 | 3.8 | 27.5 | 30.0 | 100.0 | 388.6 |
| $s=n_{\text {on }}-\hat{\mu}_{\mathrm{b}}$ | 3.0 | 4.7 | 5.2 | 13.2 | 22.5 | 37 | 100 | 134 |
| $\sigma_{\mathrm{b}}$ | 0.447 | 0.3 | 0.9 | 0.6 | 3.71 | 7.75 | 31.6 | 8.1 |
| $f=\sigma_{\mathrm{b}} / \hat{\mu}_{\mathrm{b}}$ | 0.447 | 0.231 | 0.237 | 0.158 | 0.135 | 0.258 | 0.316 | 0.0207 |
| Reported $p$ |  | 0.003 | 0.027 | 2E-06 |  |  |  |  |
| Reported Z |  | 2.7 | 1.9 | 4.6 |  |  |  | 5.9 |
| See conclusion |  |  |  |  |  |  |  |  |
| $Z_{\mathrm{Bi}}=Z_{\Gamma}$ binomial | 1.66 | 2.63 | 1.82 | 4.46 | 2.93 | 2.89 | 2.20 | 5.93 |
| $Z_{\mathrm{N}}$ Bayes Gaussian | 1.88 | 2.71 | 1.94 | 4.55 | 3.08 | 3.44 | 2.90 | 5.93 |
| $Z_{\text {PL }}$ profile likelihood | 1.95 | 2.81 | 1.99 | 4.57 | 3.02 | 3.04 | 2.38 | 5.93 |
| $Z_{\text {ZR }}$ variance stabilization | 1.93 | 2.66 | 1.98 | 4.22 | 3.00 | 3.07 | 2.39 | 5.86 |
| Not recommended |  |  |  |  |  |  |  |  |
| $Z_{\text {BiN }}=s / \sqrt{n_{\text {tot }} / \tau}$ | 2.24 | 3.59 | 2.17 | 5.67 | 3.11 | 2.89 | 2.18 | 6.16 |
| $Z_{\mathrm{nn}}=s / \sqrt{n_{\text {on }}+n_{\text {off }} / \tau^{2}}$ | 1.46 | 1.90 | 1.66 | 3.17 | 2.82 | 3.28 | 2.89 | 5.54 |
| $Z_{\text {ssb }}=s / \sqrt{\mu_{\mathrm{b}}+s}$ | 1.50 | 1.92 | 1.73 | 3.20 | 3.18 | 4.52 | 7.07 | 5.88 |
| $Z_{\mathrm{b} 0}=s / \sqrt{n_{\text {off }}(1+\tau) / \tau^{2}}$ | 2.74 | 3.99 | 2.42 | 6.47 | 3.50 | 3.90 | 3.02 | 6.31 |
| Ignore $\sigma_{\mathrm{b}}$ |  |  |  |  |  |  |  |  |
| $Z_{\mathrm{P}}$ Poisson: ignore $\sigma_{\mathrm{b}}$ | 2.08 | 2.84 | 2.14 | 4.87 | 3.80 | 5.76 | 8.76 | 6.44 |
| $Z_{\text {sb }}=s / \sqrt{\mu_{\mathrm{b}}}$ | 3.00 | 4.12 | 2.67 | 6.77 | 4.29 | 6.76 | 10.00 | 6.82 |
| Unsuccessful ad hockery |  |  |  |  |  |  |  |  |
| Poisson: $\mu_{\mathrm{b}} \rightarrow \hat{\mu}_{\mathrm{b}}+\sigma_{\mathrm{b}}$ | 1.56 | 2.51 | 1.64 | 4.47 | 3.04 | 4.24 | 5.51 | 6.01 |
| $s / \sqrt{\mu_{\mathrm{b}}+\sigma_{\mathrm{b}}}$ | 2.49 | 3.72 | 2.40 | 6.29 | 4.03 | 6.02 | 8.72 | 6.75 |






## What did we learn?

## Shapes of tails matter at 3-5 sigma

ZBi: no undercoverage; can overcover for small N Recommended
ZPL quite reasonable behavior (despite Wilks); pretty fast to calculate
ZN undercovers, worse at high Z (Cranmer)
$\mathrm{S} / \sqrt{\mathrm{B}}$ and $\mathrm{S} / \sqrt{ }(\mathrm{S}+\mathrm{B})$ : just don't: tails are wrong
Small $\tau$ is hard to get right

## Summary

Should use Binomial Test for small N, large Zmin
Good Frequentist Properties
smallest N, overcovers a bit
numerically, more work than ZpL
Binomial Test and L. Ratio have roots in Hyp Testing
For high and moderate N, Zpl Likelihood Ratio Good
Not so much for low $N$ or negative
Most wrong formulae overestimate significance
$\mathrm{S} / \sqrt{ } \mathrm{B}$ is way too optimistic-ignores uncertainty in B
You MUST check properties
Zn has coverage problems at large Zmin

## References

Li \& Ma Astroph. Journ. 272 (1983) 314-324 Zhang \& Ramsden Experimental Astronomy 1 (1990) 145-163 Fraser Journ. Am. Stat. Soc. 86 (1990) 258-265 Alexandreas et. al. Nuc. Inst. \& Meth. A328 (1993) 570-577

Gelman et. al., Bayesian Data Analysis, Chapman \& Hall (1998) (predictive p -value terminology)

